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Fano hypersurfaces and Calabi-Yau supermanifolds

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ABSTRACT: In this paper, we study the geometrical interpretations associated with Sethi's proposed general correspondence between $\mathcal{N}=2$ Landau-Ginzburg orbifolds with integral \hat{c} and $\mathcal{N}=2$ nonlinear sigma models. We focus on the supervarieties associated with $\hat{c}=3$ Gepner models. In the process, we test a conjecture regarding the superdimension of the singular locus of these supervarieties. The supervarieties are defined by a hypersurface $\widetilde{W}=0$ in a weighted superprojective space and have vanishing super-first Chern class. Here, \widetilde{W} is the modified superpotential obtained by adding as necessary to the Gepner superpotential a boson mass term and/or fermion bilinears so that the superdimension of the supervariety is equal to \hat{c} . When Sethi's proposal calls for adding fermion bilinears, setting the bosonic part of \widetilde{W} (denoted by \widetilde{W}_{bos}) equal to zero defines a Fano hypersurface embedded in a weighted projective space. In this case, if the Newton polytope of \widetilde{W}_{bos} admits a nef partition, then the Landau-Ginzburg orbifold can be given a geometrical interpretation as a nonlinear sigma model on a complete intersection Calabi-Yau manifold. The complete intersection Calabi-Yau manifold should be equivalent to the Calabi-Yau supermanifold prescribed by Sethi's proposal.

Keywords: Conformal Field Models in String Theory, String Duality, Sigma Models

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1 Introduction

Contents

Mirror symmetry [1] is a duality between string theories propagating on distinct but mirror target spaces. Consider a string theory compactified on a Calabi-Yau manifold X which is related by mirror symmetry to a string theory compactified on a Calabi-Yau manifold Y. The mirror map relates the Hodge numbers of X and Y by $h_X^{p,q} = h_Y^{\mathcal{D}-p,q}$, where \mathcal{D} is the complex dimension of X and Y. Thus, the mirror map identifies the complex structure moduli space of X with the Kähler moduli space of Y and vice-versa.

A rigid Calabi-Yau manifold has no complex structure moduli. The mirror of such a manifold has no Kähler moduli and hence cannot be a Kähler manifold in the conventional sense. Thus, Calabi-Yau manifolds cannot be the most general geometrical framework for understanding mirror symmetry. The first progress towards generalizing this framework came when Schimmrigk [2] suggested that higher-dimensional Fano varieties could provide mirrors for rigid Calabi-Yau manifolds. The name "generalized Calabi-Yau" was introduced by Candelas et al. [3] for these mirror manifolds. Later progress came when Sethi [4] proposed a general correspondence between $\mathcal{N}=2$ Landau-Ginzburg orbifolds with integral $\hat{c} \equiv c/3$ (where c is the central charge) and $\mathcal{N}=2$ nonlinear sigma models. Here, the target space of the nonlinear sigma model is either a Calabi-Yau manifold or a Calabi-Yau supermanifold. Using this proposal, Sethi argued that the mirror of a rigid Calabi-Yau manifold is a Calabi-Yau supermanifold and hence mirror symmetry should be viewed as a relation among Calabi-Yau manifolds and Calabi-Yau supermanifolds alike. The bodies of the supermanifolds are the Fano varieties mentioned above. Witten [5] described $\mathcal{N}=2$ nonlinear sigma models on Calabi-Yau manifolds and $\mathcal{N}=2$ Landau-Ginzburg orbifolds as being different phases of $\mathcal{N}=2$ gauged linear sigma models. Sethi's work and [6] have led others [7] to study $\mathcal{N}=2$ gauged linear sigma models which have a phase described by an $\mathcal{N}=2$ nonlinear sigma model on a Calabi-Yau supermanifold. The gauged linear

sigma model framework allows the argument establishing a correspondence between the nonlinear sigma model and the Landau-Ginzburg orbifold to be made more robust.

A Calabi-Yau supermanifold M obtained from Sethi's proposed correspondence would be realized by resolving the singularities of a supervariety \mathcal{M} embedded in a weighted superprojective space

$$WSP^{(n|2m)} \equiv WSP(n_{z_1}, \dots, n_{z_{n+1}} | n_{\eta_1}, \dots, n_{\eta_{2m}}).$$
(1.1)

Here, \mathcal{M} is defined by the zero locus of a transverse,¹ quasihomogeneous superpotential $\widetilde{W} = \widetilde{W}(z_{\mu}; \eta_{\alpha})$, where

$$z_{\mu} \simeq \lambda^{n_{z_{\mu}}} z_{\mu} \ (\mu = 1, \dots, n+1), \quad \eta_{\alpha} \simeq \lambda^{n_{\eta_{\alpha}}} \eta_{\alpha} \ (\alpha = 1, \dots, 2m)$$
 (1.2)

are homogeneous bosonic and fermionic coordinates of weights $n_{z_{\mu}}$ and $n_{\eta_{\alpha}}$, respectively. Since \widetilde{W} is quasihomogeneous, it satisfies

$$\widetilde{W}(\lambda^{n_{z_{\mu}}} z_{\mu}; \lambda^{n_{\eta_{\alpha}}} \eta_{\alpha}) = \lambda^{d} \widetilde{W}(z_{\mu}; \eta_{\alpha}), \qquad (1.3)$$

where d is the degree of quasihomogeneity. The superpotential \widetilde{W} is obtained from the transverse, quasihomogeneous superpotential $W=W(\Phi_a)$ of an $\mathcal{N}=2$ Landau-Ginzburg model with integral \hat{c} by truncating each chiral superfield Φ_a $(a=1,\ldots,N)$ to its lowest bosonic component ϕ_a , setting $\phi_a=z_a$, and then adding boson mass terms $z_{N+1}^2+\cdots+z_{n+1}^2$ and/or fermion bilinears $\eta_1\eta_2+\cdots+\eta_{2m-1}\eta_{2m}$ to W so that

$$\widetilde{\mathcal{D}} \equiv (n+1) - 2m - 2 = \hat{c}. \tag{1.4}$$

Here, $\widetilde{\mathcal{D}}$ is the *superdimension* of the Calabi-Yau supermanifold. The condition (1.4) allows a change of variables with constant Jacobian to be made such that one of the new variables appears only linearly in the modified superpotential \widetilde{W} . The change of variables is not one-to-one, so the modified theory must be orbifolded by the diagonal subgroup of its phase symmetries. Integrating the linear new variable out of the path integral for the modified action as a Lagrange multiplier yields a super-delta function constraint which corresponds to \mathcal{M} having vanishing super-first Chern class.

When $m \neq 0$, setting the bosonic part of \widetilde{W} (denoted by \widetilde{W}_{bos}) equal to zero defines a Fano hypersurface \mathcal{F} embedded in a weighted projective space

$$\mathbf{WP}^n \equiv \mathbf{WP}(n_{z_1}, \dots, n_{z_{n+1}}). \tag{1.5}$$

In this case, if the Newton polytope Δ corresponding to \widetilde{W}_{bos} admits a nef partition $\Delta = \Delta_1 + \cdots + \Delta_r$, then the Landau-Ginzburg orbifold can be given a geometrical interpretation as a nonlinear sigma model on a complete intersection Calabi-Yau manifold defined by equations $f_i = 0$ (i = 1, ..., r) [8]. Here, Δ_i is the Newton polytope of f_i . The complete intersection Calabi-Yau manifold should be equivalent (in the sense of [6]) to the Calabi-Yau supermanifold prescribed by Sethi's proposed correspondence. When m = 0, the

¹Transverse \widetilde{W} means that $\widetilde{W} = 0$ and $d\widetilde{W} = 0$ have no common solution except at the origin.

constraint $\widetilde{W} = 0$ defines a Calabi-Yau variety \mathcal{X} embedded in \mathbf{WP}^n . Resolving \mathcal{X} would yield a Calabi-Yau manifold X of complex dimension

$$\mathcal{D} \equiv (n+1) - 2 = \hat{c}. \tag{1.6}$$

The complex dimension of the singular locus of \mathcal{X} satisfies [9]

$$0 \le \dim(\operatorname{Sing}(\mathcal{X})) \le \mathcal{D} - 2$$
, $\operatorname{Sing}(\mathcal{X}) = \mathcal{X} \cap \operatorname{Sing}(\mathbf{WP}^n)$. (1.7)

To obtain a Calabi-Yau supermanifold M by resolving the singularities of \mathcal{M} , one might infer from the discussion in [4] that the superdimension of the singular locus of \mathcal{M} must satisfy

$$\operatorname{sdim}(\operatorname{Sing}(\mathcal{M})) \leq \widetilde{\mathcal{D}} - 1, \quad \operatorname{Sing}(\mathcal{M}) = \mathcal{M} \cap \operatorname{Sing}\left(\mathbf{WSP}^{(n|2m)}\right).$$
 (1.8)

The result (1.7) and the equivalence discussed in [6] suggests the following stronger conjecture:

Conjecture 1.1 To obtain a Calabi-Yau supermanifold M by resolving the singularities of \mathcal{M} , the superdimension of the singular locus of \mathcal{M} must satisfy

$$\mathrm{sdim}(\mathrm{Sing}(\mathcal{M})) \leq \widetilde{\mathcal{D}} - 2\,, \quad \mathrm{Sing}(\mathcal{M}) = \mathcal{M} \cap \mathrm{Sing}\left(\mathbf{WSP}^{(n|2m)}\right)\,.$$

In this paper, we will test the above conjecture for $\widetilde{\mathcal{D}}=3$. This will be achieved by studying the geometrical interpretations prescribed by Sethi's proposed correspondence for Gepner models [10] with $\hat{c}=3$. Since \widetilde{W} is quasihomogeneous of degree d, the weights of the fermions in each fermion bilinear $\eta_{2k-1}\eta_{2k}$ $(k=1,\ldots,m)$ must satisfy

$$n_{\eta_{2k-1}} + n_{\eta_{2k}} = d. (1.9)$$

Requiring either the singular locus constraint (1.8) or conjecture 1.1 to hold further restricts the fermionic weights. We have written a computer program which allows these restrictions to be implemented. In principle, one could determine the fermionic weights by requiring agreement between the Hodge diamond of the Landau-Ginzburg orbifold and the Hodge diamond of \mathcal{M} . The former Hodge diamond can be computed using the techniques of [11] whereas insight into the structure of the latter Hodge diamond can be obtained from the heuristic approach of [4] based on orbifold considerations [12]. We will compare the fermionic weights obtained in this way with those obtained from our computer program.

This paper is organized as follows: In section 2, we discuss the application of Sethi's proposed Landau-Ginzburg orbifold/nonlinear sigma model correspondence to Gepner models. In section 3, we describe how the singular locus constraint (1.8) restricts the fermionic weights and work through an example. It is a trivial step to replace (1.8) with conjecture 1.1 in our computer program. The analysis of section 4 compares the fermionic weights obtained from our computer program with those obtained from the cohomological approach described in the previous paragraph. Several examples are included which highlight the similarities and differences. Concluding remarks are given in section 5. Finally, in the appendix, we tabulate the families of hypersurfaces associated with our supervarieties.

Here, the fermionic weights are determined with our computer program by requiring (1.8) and (1.9) to be satisfied. For each hypersurface family, the Hodge numbers $h^{1,1}$ and $h^{2,1}$ and the Euler number of the associated Landau-Ginzburg orbifold are given. When the Newton polytope corresponding to \widetilde{W}_{bos} admits a nef partition, this is indicated. We also indicate when the Newton polytope of \widetilde{W}_{bos} is nonreflexive Gorenstein.

2 Gepner/NLSM correspondence

The worldsheet action for an $\mathcal{N}=2$ Landau-Ginzburg model is [13]

$$S = \int d^2 \mathbf{z} \, d^4 \theta \, K \left(\Phi_a, \overline{\Phi}_a \right) + \left(\int d^2 \mathbf{z} \, d^2 \theta \, W(\Phi_a) + c.c. \right), \tag{2.1}$$

Here, the integral involving the Kähler potential K is called the D-term, the integral involving the superpotential W is called the F-term, and Φ_a ($a=1,\ldots,N$) are chiral superfields. The D-term contains only irrelevant operators whereas the F-term contains relevant operators. Thus, the superpotential defines a universality class under renormalization group flow. Requiring the superpotential to be transverse and quasihomogeneous is believed to ensure the existence of a unique, nontrivial IR fixed point which is conformally invariant. At this fixed point, the action (2.1) provides a Lagrangian description of an $\mathcal{N}=2$ minimal model [13–15] with

$$\hat{c} = 2 \sum_{a=1}^{N} \left(\frac{1}{2} - q_{\Phi_a} \right) , \quad q_{\Phi_a} \equiv n_{\Phi_a} / d .$$
 (2.2)

The 10,839 transverse, quasihomogeneous Landau-Ginzburg superpotentials corresponding to $\mathcal{N}=2$ superconformal theories with $\hat{c}=3$ were classified in [16]. A subset of these correspond to Gepner models with $\hat{c}=3$. A Gepner model [10] is a string model constructed as an orbifold of a tensor product of $\mathcal{N}=2$ minimal models. The central charge of the ith $\mathcal{N}=2$ minimal model in the tensor product is

$$c_i = \frac{3k_i}{k_i + 2} \quad (k_i = 1, 2, \dots),$$
 (2.3)

where k_i is the level of the $\mathcal{N}=2$ superconformal algebra [17]. To obtain an anomaly-free compactification to D spacetime dimensions (D<10, even), the internal contribution to the central charge must be

$$c = \sum_{i} c_i = \frac{3}{2} (10 - D). \tag{2.4}$$

The work of [18, 19] associated Calabi-Yau manifolds to a large class of Gepner models. Sethi's proposal [4] prescribes a geometrical interpretation for *all* Gepner models. This prescription is as follows:

- 1. Start by associating superpotential terms [14]
 - $W_i = x_i^{k_i+2}$ to A-models of level k_i ,
 - $W_i = x_i^{\frac{k_i}{2}+1} + x_i y_i^2$ to *D*-models of level k_i (even),

• $W_i = x_i^3 + y_i^4$, $W_i = x_i^3 + x_i y_i^3$, and $W_i = x_i^3 + y_i^5$ to E_{6} -, E_{7} -, and E_{8} -models of level $k_i = 10, 16, 28$, respectively.

The tensor product of r subtheories yields a transverse, quasihomogeneous Gepner superpotential

$$W = \sum_{i=1}^{r} W_i = \sum_{i=1}^{r} \left(x_i^{l_{x_i}} + x_i^{\bar{l}_{x_i}} y_i^{l_{y_i}} \right). \tag{2.5}$$

For A-models, $y_i = 0$ and $l_{x_i} = k_i + 2$. For D-models, $l_{x_i} = k_i/2 + 1$, $\bar{l}_{x_i} = 1$, and $l_{y_i} = 2$. For E_{6^-} , E_{7^-} , and E_{8^-} models, $l_{x_i} = 3, 3, 3$, $\bar{l}_{x_i} = 0, 1, 0$, and $l_{y_i} = 4, 3, 5$, respectively. In all cases,

$$n_{x_i}l_{x_i} = d. (2.6)$$

Additionally, for D- and E-models.

$$n_{x_i}\bar{l}_{x_i} + n_{y_i}l_{y_i} = d. (2.7)$$

The x_i and nonzero y_i (i = 1, ..., r) are identified with the z_a (a = 1, ..., N) described in the introduction according to the convention $z_1 = x_1$,

$$z_2 = \begin{cases} y_1 & (y_1 \neq 0) \\ x_2 & (y_1 = 0), \end{cases}$$

and so on.

2. Add as necessary to W a single² boson mass term z^2 and/or fermion bilinears $\eta_1\eta_2 + \cdots + \eta_{2m-1}\eta_{2m}$ so that

$$(n+1) - 2m - 2 = \hat{c}. (2.8)$$

We thus obtain the modified superpotential

$$\widetilde{W} = \begin{cases} W & (\hat{c} = N - 2) \\ W + z^2 & (\hat{c} > N - 2) \\ W + \sum_{k=1}^{m} \eta_{2k-1} \eta_{2k} & (\hat{c} < N - 2, \sum_{a=1}^{N} q_{z_a} \text{integral}) \\ W + z^2 + \sum_{k=1}^{m} \eta_{2k-1} \eta_{2k} & (\hat{c} < N - 2, \sum_{a=1}^{N} q_{z_a} \text{half-integral}) . \end{cases}$$
(2.9)

The added fields have no effect on the chiral ring or the conformal fixed point to which the theory flows. The condition (2.8) allows a change of variables with constant Jacobian to

- For $\hat{c} = N 2$, no fields need to be added.
- For $\hat{c} > N-2$, boson mass terms $z_{N+1}^2 + \cdots + z_{n+1}^2$ are required.
- For $\hat{c} < N-2$, the condition that \hat{c} be integral implies that the sum of the charges $\sum_{a=1}^{N} q_{z_a}$ can be either integral or half-integral. The former requires adding fermion bilinears $\eta_1 \eta_2 + \cdots + \eta_{2m-1} \eta_{2m}$ whereas the latter requires adding a single boson mass term z_{n+1}^2 and fermion bilinears $\eta_1 \eta_2 + \cdots + \eta_{2m-1} \eta_{2m}$.

For Gepner models, \hat{c} is never greater than N-2 by more than 1. Thus, it is never necessary to add more than one boson mass term to the Gepner superpotential.

²There are three possibilities for an arbitrary Landau-Ginzburg model with integral \hat{c} :

be made such that one of the new variables appears only linearly in the modified superpotential. The change of variables is not one-to-one, so the modified theory must orbifolded by the diagonal subgroup of its phase symmetries. When m=0 ($m\neq 0$), integrating the linear new variable out of the path integral for the modified action as a Lagrange multiplier yields a (super-)delta function constraint which corresponds to the bosonic variety $\mathcal X$ (supervariety $\mathcal M$) defined by $\widetilde W=0$ having vanishing (super-)first Chern class. The first Chern class of $\mathcal X$ vanishes when

$$\sum_{\mu=1}^{n+1} n_{z_{\mu}} - d = 0, \qquad (2.10)$$

whereas the super-first Chern class of \mathcal{M} vanishes when

$$\sum_{\mu=1}^{n+1} n_{z_{\mu}} - \sum_{\alpha=1}^{2m} n_{\eta_{\alpha}} - d = 0.$$
 (2.11)

3 Fermionic weights and the singular locus constraint

As described in the introduction, we require the fermionic weights to be consistent with the singular locus constraint (1.8) and the quasihomogeneity constraint (1.9). The restriction placed on the fermionic weights by the latter constraint is obvious. Let us now explain what consistency with the former constraint means.

The supervariety \mathcal{M} defined by the hypersurface $\widetilde{W} = 0$ has a \mathbf{Z}_p fixed point set under the weighted projective identification (1.2) if and only if the following conditions are both satisfied:

- 1. The index set $B_p \equiv \{\mu | \lambda_p^{n_{z\mu}} = 1\}$ is nonempty, where $\lambda_p \equiv e^{2\pi i/p}$.
- 2. The quantity

$$\mathcal{D}_p = \begin{cases} |B_p| - 2 & (d/p \in \mathbf{Z}) \\ |B_p| - 1 & (d/p \notin \mathbf{Z}). \end{cases}$$
(3.1)

satisfies $\mathcal{D}_p \geq 0$ if the index set $F_p \equiv \{\alpha | \lambda_p^{n_{\eta_\alpha}} = 1\}$ is empty.

There are no purely fermionic \mathbf{Z}_p fixed point sets because the body of \mathcal{M} is embedded in \mathbf{WP}^n . The superdimension of the \mathbf{Z}_p fixed point sets which do exist is given by

$$\widetilde{\mathcal{D}}_{p} = \begin{cases} |B_{p}| - |F_{p}| - 2 & (d/p \in \mathbf{Z}) \\ |B_{p}| - |F_{p}| - 1 & (d/p \notin \mathbf{Z}). \end{cases}$$
(3.2)

In the case $d/p \in \mathbf{Z}$, the -2 arises because the weighted projective identification and the hypersurface equation each reduce the superdimension by 1. In the case $d/p \notin \mathbf{Z}$, the hypersurface equation is identically satisfied and hence does not reduce the superdimension. Consistency with the singular locus constraint (1.8) means that

$$\widetilde{\mathcal{D}}_p \le \hat{c} - 1 \quad \forall p \,, \tag{3.3}$$

where we have used (1.4). We see from (3.2) that \mathbf{Z}_p fixed point sets with $|B_p| \leq |F_p|$ have negative superdimension and hence are consistent with (1.8). Thus, when checking for consistency with (1.8), we can focus our attention on \mathbf{Z}_p fixed point sets with $|B_p| > |F_p|$. Note that the bosonic part of such a \mathbf{Z}_p fixed point set has complex dimension given by (3.1).

To illustrate the above, let us consider a concrete example.

Example 3.1 Consider a Gepner model with level/invariant structure 10_D 10_D 1_A 1_A 1_A . Following the procedure described in section 2, we obtain the quasihomogeneous degree d = 12 modified superpotential

$$\widetilde{W} = \sum_{i=1}^{2} (x_i^6 + x_i y_i^2) + \sum_{i=3}^{6} x_i^3 + z^2 + \eta_1 \eta_2 + \eta_3 \eta_4.$$

The hypersurface $\widetilde{W} = 0$ defines a supervariety \mathcal{M} embedded in

WSP
$$(2, 5, 2, 5, 4, 4, 4, 4, 6 | n_{\eta_1}, n_{\eta_2}, n_{\eta_3}, n_{\eta_4})$$
.

We will denote the family of quasihomogeneous degree d = 12 hypersurfaces embedded in this weighted superprojective space by

WSP
$$(2, 5, 2, 5, 4, 4, 4, 4, 6 | n_{\eta_1}, n_{\eta_2}, n_{\eta_3}, n_{\eta_4})$$
[12].

A hypersurface in this family would also be obtained from a Gepner model with level and invariant structure $4_D \ 4_D \ 10_D \ 0r \ 4_D \ 10_D \ 10_D \ 1_A \ 1_A$.

According to (1.9), the fermions in each bilinear can take on the values (modulo a relabelling of the fermions)

$$(n_{\eta_{2k-1}}, n_{\eta_{2k}}) \in \{(1, 11), (2, 10), (3, 9), (4, 8), (5, 7), (6, 6)\}. \tag{3.4}$$

To further constrain the fermionic weights, we now consider the \mathbf{Z}_p (p=2,4,5) fixed point sets. First, we determine the complex dimension of the bosonic parts of these fixed point sets:

p = 2: The bosonic part of the \mathbb{Z}_2 fixed point set is

$$\sum_{i=1}^{2} (x_i^6 + x_i y_i^2) + \sum_{i=3}^{6} x_i^3 + z^2 = 0, \quad y_1 = y_2 = 0.$$

There are $|B_2| = 7$ bosons $(x_i \ (i = 1, ..., 6) \text{ and } z)$ in this fixed point set. Since $d/p = 6 \in \mathbb{Z}$, (3.1) gives

$$\mathcal{D}_2 = |B_2| - 2 = 5. (3.5)$$

p = 4: The bosonic part of the \mathbb{Z}_4 fixed point set is

$$\sum_{i=3}^{6} x_i^3 = 0, \quad x_i = y_i = 0 \quad (i = 1, 2), \quad z = 0.$$

j	$(n_{\eta_1}, n_{\eta_2}, n_{\eta_3}, n_{\eta_4})_j$	$\widetilde{\mathcal{D}}_2^{(j)}$	$\widetilde{\mathcal{D}}_4^{(j)}$	$\widetilde{\mathcal{D}}_{5}^{(j)}$
1	(1,11,1,11)	5	2	1
2	(1,11,2,10)	3	2	0
3	(1,11,3,9)	5	2	1
4	(1,11,4,8)	3	0	1
5	(1,11,5,7)	5	2	0
6	(1,11,6,6)	3	2	1
7	(2,10,2,10)	1	2	-1
8	(2,10,3,9)	3	2	0
9	(2,10,4,8)	1	0	0
10	(2,10,5,7)	3	2	-1
11	(2,10,6,6)	1	2	0
12	(3,9,3,9)	5	2	1
13	(3,9,4,8)	3	0	1
14	(3,9,5,7)	5	2	0
15	(3,9,6,6)	3	2	1
16	(4,8,4,8)	1	-2	1
17	(4,8,5,7)	3	0	0
18	(4,8,6,6)	1	0	1
19	(5,7,5,7)	5	2	-1
20	(5,7,6,6)	3	2	0
21	(6,6,6,6)	1	2	1

Table 1. Computation of $\widetilde{\mathcal{D}}_p^{(j)} = \mathcal{D}_p - |F_p^{(j)}| \ (p = 2, 4, 5).$

There are $|B_4|=4$ bosons $(x_i \ (i=3,\ldots,6))$ in this fixed point set. Since $d/p=3\in \mathbb{Z}$, (3.1) gives

$$\mathcal{D}_4 = |B_4| - 2 = 2. \tag{3.6}$$

p = 5: The bosonic part of the \mathbb{Z}_5 fixed point set is

$$\sum_{i=1}^{2} (x_i^6 + x_i y_i^2) = 0, \quad x_i = 0 \quad (i = 1, \dots, 6), \quad z = 0.$$

There are $|B_5| = 2$ bosons $(y_1 \text{ and } y_2)$ in this fixed point set. Since $d/p = 12/5 \notin \mathbb{Z}$, (3.1) gives

$$\mathcal{D}_5 = |B_5| - 1 = 1. (3.7)$$

Next, let $\widetilde{\mathcal{D}}_p^{(j)} = \mathcal{D}_p - |F_p^{(j)}|$ be the superdimension of the \mathbf{Z}_p fixed point set for the jth distinct (modulo relabelling) fermionic weight combination $(n_{\eta_1}, n_{\eta_2}, n_{\eta_3}, n_{\eta_4})_j$ consistent with (1.9). For m = 2 fermion bilinears, there are $\frac{1}{2} \left[\frac{d}{2} \right] \left(\left[\frac{d}{2} \right] + 1 \right)$ such combinations, where [x] denotes the integer part of x. Thus, for this example, $j = 1, \ldots, 21$. The $\widetilde{\mathcal{D}}_p^{(j)}$ are given in table 1.

Finally, exclude all fermionic weight combinations which do not satisfy

$$\widetilde{\mathcal{D}}_p^{(j)} \le \hat{c} - 1 = 2 \ \forall p \,.$$

This leaves the fermionic weight combinations

$$(2,10,2,10)$$
, $(2,10,4,8)$, $(2,10,6,6)$, $(4,8,4,8)$, $(4,8,6,6)$, $(6,6,6,6)$.

We can use the parameters k and l defined in the appendix to express these combinations in the compact form (2k, 12-2k, 2l, 12-2l). Here, $k=1,\ldots,3$ and $l=1,\ldots,3$. It is understood that repeated fermionic weight combinations generated with this notation are counted only once. In this notation, the hypersurface $\widetilde{W}=0$ defines a supervariety \mathcal{M} embedded in $\mathbf{WSP}(2,5,2,5,4,4,4,4,6|2k,12-2k,2l,12-2l)$ and is a member of hypersurface family 250 of the appendix:

$$WSP(2, 5, 2, 5, 4, 4, 4, 4, 6 | 2k, 12 - 2k, 2l, 12 - 2l)[12]$$
.

4 Analysis

We have written a computer program which takes as input data which encodes the superpotential W for $\hat{c}=3$ Gepner models. Next, the program determines the modified superpotential \widetilde{W} as explained in section 2. When \widetilde{W} depends on five bosonic fields and no fermionic fields, the output of the program is the hypersurface family

$$\mathbf{WP}(n_{z_1},\ldots,n_{z_{n+1}})[d]$$

corresponding to the hypersurface $\widetilde{W} = 0$ which defines a bosonic variety \mathcal{X} embedded in $\mathbf{WP}(n_{z_1}, \dots, n_{z_{n+1}})$. For the remaining cases, the output is the hypersurface family

WSP
$$(n_{z_1}, \dots, n_{z_{n+1}} | n_{\eta_1}, \dots, n_{\eta_{2m}})[d]$$

corresponding to the hypersurface $\widetilde{W}=0$ which defines a supervariety \mathcal{M} embedded in $\mathbf{WSP}(n_{z_1},\ldots,n_{z_{n+1}}|n_{\eta_1},\ldots,n_{\eta_{2m}})$. The fermionic weights are determined by requiring the singular locus constraint (1.8) and the quasihomogeneity constraint (1.9) to be satisfied. In this manner, we obtain at least one fermionic weight solution for each $\hat{c}=3$ Gepner model corresponding to a supervariety \mathcal{M} through Sethi's proposed correspondence. The 254 hypersurface families associated with these supervarieties are tabulated in the appendix. When we replace the singular locus constraint (1.8) with conjecture 1.1, this results in models corresponding to the hypersurface families 50, 94, 95, 121, 125, and 229 of the appendix having no solution for the fermionic weights.

In principle, the singularities which do arise could be determined by identifying the fermionic weights which yield agreement between the Hodge diamond of the Landau-Ginzburg orbifold and the Hodge diamond of \mathcal{M} . The Landau-Ginzburg orbifold Hodge diamond can be computed using the techniques of [11]. Such calculations can be done quickly with the software package PALP [20]. Unfortunately, at present, there is no supercohomology theory which allows the Hodge diamond of \mathcal{M} to be computed.

The Hodge diamond of \mathcal{X} can be computed using the orbifold techniques of [12]. This is possible because, with a change of coordinates

$$z_{\mu} = (\zeta_{\mu})^{n_{z_{\mu}}}, \qquad (4.1)$$

the hypersurface $\mathbf{WP}(n_{z_1}, \dots, n_{z_{n+1}})[d]$ which defines \mathcal{X} can be written as an orbifold of a hypersurface in a homogeneous projective space \mathbf{P}^n , i.e.

$$\mathbf{WP}(n_{z_1},\dots,n_{z_{n+1}})[d] = \frac{\mathbf{P}^n[d]}{\mathbf{Z}_{n_{z_1}} \times \dots \times \mathbf{Z}_{n_{z_{n+1}}}}.$$
(4.2)

In contrast, the hypersurface $\mathbf{WSP}(n_{z_1}, \dots, n_{z_{n+1}} | n_{\eta_1}, \dots, n_{\eta_{2m}})[d]$ which defines \mathcal{M} cannot be written as an orbifold of a hypersurface in homogeneous superprojective space $\mathbf{SP}^{(n|2m)}$. This is because the analogue of (4.1) does not make sense for Grassmann coordinates. Nevertheless, as described in [4], we can use orbifold considerations to gain insight into the structure of the Hodge diamond of \mathcal{M} . In the following examples, we will use this heuristic reasoning and compare the resulting fermionic weights with those obtained from the singular locus constraint (1.8) and conjecture 1.1.

Example 4.1 A hypersurface in the family $WSP(1,1,2,2,2,2,2|n_{\eta_1},n_{\eta_2})[6]$ can be obtained from a Gepner model with any of the following level/invariant structures:

$$4_D \ 4_D \ 4_A \ 4_A \ 1_A$$
, $4_D \ 4_A \ 4_A \ 1_A \ 1_A \ 1_A$, $4_A \ 4_A \ 1_A \ 1_A \ 1_A \ 1_A$.

For definiteness, we will focus on the last of these. Following the procedure described in section 2, we obtain the modified superpotential

$$\widetilde{W} = x_1^6 + x_2^6 + x_3^3 + x_4^3 + x_5^3 + x_6^3 + x_7^3 + \eta_1 \eta_2$$
.

This example was discussed in [4]. Here, we simply note that cohomology considerations, conjecture 1.1, and, as indicated by hypersurface family 179 of the appendix, the singular locus constraint (1.8) all yield the same unique result $(n_{\eta_1}, n_{\eta_2}) = (2, 4)$.

Example 4.2 A hypersurface in the family $WSP(1, 1, 4, 4, 4, 4, 6|n_{\eta_1}, n_{\eta_2})[12]$ can be obtained from a Gepner model with any of the following level/invariant structures:

$$4_D \ 10_A \ 4_A \ 2_A \ 1_A$$
, $10_A \ 4_A \ 2_A \ 1_A \ 1_A \ 1_A$, $4_D \ 4_D \ 10_A \ 10_A$, $4_D \ 10_A \ 1_A \ 1_A$, $10_A \ 10_A \ 1_A \ 1_A \ 1_A \ 1_A$.

For definiteness, we will focus on the last of these. Following the procedure described in section 2, we obtain the modified superpotential

$$\widetilde{W} = x_1^{12} + x_2^{12} + x_3^{12} + x_3^{3} + x_4^{3} + x_5^{3} + x_6^{3} + z^{2} + \eta_1 \eta_2$$
.

Employing the heuristic reasoning of [4], we find for $(n_{\eta_1}, n_{\eta_2}) = (4, 8)$ that the Hodge diamond of \mathcal{M} is

The first term on the right-hand side of (4.3) is the contribution arising from the untwisted sector. The second term includes the contribution of the fixed point set associated with the third twisted sector (the upper 6) and the fixed point set associated with the ninth twisted sector (the lower 6). Finally, the last term arises from the identity and volume forms of the fixed point set associated with the sixth twisted sector. Our result (4.3) agrees with the Hodge diamond of the associated Landau-Ginzburg orbifold. For all other fermionic weight assignments consistent with (1.9), we do not obtain this agreement. Thus, in this example, using these heuristic arguments, we obtain a unique result

$$(n_{\eta_1}, n_{\eta_2}) = (4, 8)$$
.

We note that this result agrees with what would be obtained from conjecture 1.1. In contrast, as indicated by hypersurface family 10 of the appendix, the singular locus constraint (1.8) allows

$$(n_{\eta_1}, n_{\eta_2}) \in \{(2, 10), (4, 8), (6, 6)\}.$$

Example 4.3 A hypersurface in the family $WSP(1,3,3,3,4,4,6|n_{\eta_1},n_{\eta_2})[12]$ can be obtained from a Gepner model with either of the following level/invariant structures:

$$4_D \ 10_A \ 2_A \ 2_A \ 2_A$$
, $10_A \ 2_A \ 2_A \ 2_A \ 1_A \ 1_A$.

For definiteness, we will focus on the last of these. Following the procedure described in section 2, we obtain the modified superpotential

$$\widetilde{W} = x_1^{12} + x_2^4 + x_3^4 + x_4^4 + x_5^3 + x_6^3 + z^2 + \eta_1 \eta_2 \,.$$

Employing the heuristic reasoning of [4], we find for $(n_{\eta_1}, n_{\eta_2}) = (3, 9)$ that the Hodge diamond of \mathcal{M} is

The first three terms on the right-hand side of (4.4) have the same origin as the corresponding terms in example 4.2. The fourth term includes the contribution of the fixed point set associated with the fourth twisted sector (the lower 7) and the fixed point set associated with the ninth twisted sector (the upper 7). Our result (4.4) agrees with the Hodge diamond of the associated Landau-Ginzburg orbifold. For all other fermionic weight assignments consistent with (1.9), we do not obtain this agreement. Thus, in this example, using these heuristic arguments, we obtain a unique result

$$(n_{n_1}, n_{n_2}) = (3, 9)$$
.

In contrast, conjecture 1.1 allows

$$(n_{\eta_1}, n_{\eta_2}) \in \{(3, 9), (6, 6)\}$$

and, as indicated by hypersurface family 8 of the appendix, the singular locus constraint (1.8) allows

$$(n_{\eta_1}, n_{\eta_2}) \in \{(1, 11), (2, 10), (3, 9), (4, 8), (5, 7), (6, 6)\}.$$

Example 4.4 A hypersurface in the family $WSP(2,3,1,2,2,2,4|n_{\eta_1},n_{\eta_2})[8]$ can be obtained from a Gepner model with level/invariant structure

$$6_D 6_A 2_A 2_A 2_A$$
.

Following the procedure described in section 2, we obtain the modified superpotential

$$\widetilde{W} = x_1^4 + x_1 y_1^2 + x_2^8 + x_3^4 + x_4^4 + x_5^4 + z^2 + \eta_1 \eta_2$$
.

Employing the heuristic reasoning of [4], we find for $(n_{\eta_1}, n_{\eta_2}) = (4, 4)$ that the Hodge diamond of \mathcal{M} is

The first term on the right-hand side of (4.5) is the contribution arising from the untwisted sector whereas the second term arises from the fourth twisted sector. Our result (4.5) agrees with the Hodge diamond of the associated Landau-Ginzburg orbifold. For all other fermionic weight assignments consistent with (1.9), we do not obtain this agreement. Thus, in this example, using these heuristic arguments, we obtain a unique result

$$(n_{\eta_1}, n_{\eta_2}) = (4, 4)$$
.

In contrast, conjecture 1.1 and, as indicated by hypersurface family 3 of the appendix, the singular locus constraint (1.8) both allow

$$(n_{n_1}, n_{n_2}) \in \{(2, 6), (4, 4)\}.$$

Example 4.5 Let us revisit the family of hypersurfaces discussed in example 3.1, $\mathbf{WSP}(2,5,2,5,4,4,4,4,6|n_{\eta_1},n_{\eta_2},n_{\eta_3},n_{\eta_4})[12]$. We again focus on the hypersurface obtained from a Gepner model with level/invariant structure 10_D 10_D 1_A 1_A 1_A 1_A . Proceeding as in the above examples, we find for $(n_{\eta_1},n_{\eta_2},n_{\eta_3},n_{\eta_4})=(2,10,4,8)$ that the Hodge diamond of \mathcal{M} is

The terms on the right-hand side of (4.6) originate from the same sectors as the corresponding terms in example 4.2. Our result (4.6) agrees with the Hodge diamond of the associated Landau-Ginzburg orbifold. For all other fermionic weight assignments consistent with (1.9), we do not obtain this agreement. Thus, in this example, using these heuristic arguments, we obtain

$$(n_{\eta_1}, n_{\eta_2}, n_{\eta_3}, n_{\eta_4}) \in \{(2, 10, 4, 8), (4, 8, 6, 6)\}.$$

In contrast, conjecture 1.1 allows

$$(n_{\eta_1}, n_{\eta_2}, n_{\eta_3}, n_{\eta_4}) \in \{(2, 10, 4, 8), (4, 8, 4, 8), (4, 8, 6, 6)\}$$

and the singular locus constraint (1.8) allows

$$(n_{\eta_1}, n_{\eta_2}, n_{\eta_3}, n_{\eta_4}) \in \{(2, 10, 2, 10), (2, 10, 4, 8), (2, 10, 6, 6), (4, 8, 4, 8), (4, 8, 6, 6), (6, 6, 6, 6)\}.$$

It is interesting to note that, in the above examples, the solutions obtained from the heuristic approach agree precisely with those obtained from conjecture 1.1 supplemented by the constraint

$$\widetilde{\mathcal{D}}_p \ge 0$$
 whenever $\mathcal{D}_p \ge 0$. (4.7)

Let us now consider an example in which the heuristic approach yields no solution.

Example 4.6 A hypersurface in the family $WSP(1,1,1,2,2,2,3|n_{\eta_1},n_{\eta_2})[6]$ can be obtained from a Gepner model with either of the following level/invariant structures:

$$4_D \ 4_A \ 4_A \ 4_A \ 1_A$$
, $4_A \ 4_A \ 4_A \ 1_A \ 1_A \ 1_A$.

For definiteness, we will focus on the last of these. Following the procedure described in section 2, we obtain the modified superpotential

$$\widetilde{W} = x_1^6 + x_2^6 + x_3^6 + x_4^3 + x_5^3 + x_6^3 + z^2 + \eta_1 \eta_2$$
.

Employing the heuristic reasoning of [4], we find for $(n_{\eta_1}, n_{\eta_2}) \in \{(1, 5), (3, 3)\}$ that the Hodge diamond of \mathcal{M} is

The first term on the right-hand side of (4.8) is the contribution arising from the untwisted sector. The second term includes the contribution of the fixed point set associated with the third twisted sector. For $(n_{\eta_1}, n_{\eta_2}) = (2, 4)$, we find that the only contribution to the Hodge diamond of \mathcal{M} arises from the untwisted sector. In all cases, the result disagrees with the Landau-Ginzburg orbifold Hodge diamond

Thus, in this example, using these heuristic arguments, we find no solution for the fermionic weights. In contrast, conjecture 1.1 and, as indicated by hypersurface family 2 of the appendix, the singular locus constraint (1.8) both allow $(n_{\eta_1}, n_{\eta_2}) \in \{(1, 5), (2, 4), (3, 3)\}$.

5 Conclusion

The analysis in section 4 compares the fermionic weights obtained from our computer program with those obtained by requiring agreement between the Hodge diamond of the Landau-Ginzburg orbifold and the heuristically determined Hodge diamond of the supervariety. Running the program with the singular locus constraint (1.8) yields at least one solution for each $\hat{c}=3$ Gepner model associated with a supervariety through Sethi's proposed correspondence. Conjecture 1.1 is a stronger constraint, but it still yields at least one solution for the vast majority of these models. The heuristic approach places the strongest constraint on the fermionic weights. It yields a unique solution in examples 4.1, 4.2, 4.3, and 4.4, two solutions in example 4.5, and no solution in example 4.6. In the examples we have studied, the heuristically determined solutions are a subset of the solutions obtained from conjecture 1.1. Furthermore, in these examples, when the heuristic approach yields solutions, these solutions agree precisely with those obtained from conjecture 1.1 supplemented by the constraint (4.7). Thus, something seems to be "right" about the combination of conjecture 1.1 and the constraint (4.7). A proper supercohomology theory would allow more conclusive statements to be made.

In the appendix, table 2 indicates when the Newton polytope of \widetilde{W}_{bos} admits a nef partition. In this case, the Landau-Ginzburg orbifold can be given a geometrical interpretation as a nonlinear sigma model on a complete intersection Calabi-Yau manifold. The complete intersection Calabi-Yau manifold should be equivalent to the Calabi-Yau supermanifold prescribed by Sethi's proposal. It can be shown that a reflexive Gorenstein polytope Δ admits a nef partition if and only if the reflexive Gorenstein cones σ_{Δ} and σ_{Δ}^{\vee} are both completely split [8]. In fact, the Landau-Ginzburg orbifold can be given a complete intersection Calabi-Yau manifold interpretation even when only σ_{Δ} is completely split [8]. In example 4.4, the Newton polytope of \widetilde{W}_{bos} is nonreflexive Gorenstein. It turns out that, for all of the remaining examples in section 4, the Newton polytope of \widetilde{W}_{bos} is reflexive Gorenstein but σ_{Δ} is not completely split. Thus, in these examples, the Landau-Ginzburg orbifold cannot be given a complete intersection Calabi-Yau manifold interpretation. We leave a detailed investigation of the cases in which only σ_{Δ} or only σ_{Δ}^{\vee} is completely split to future work.

A Supervariety hypersurface families

In table 2, we list the supervariety hypersurface families associated with $\hat{c}=3$ Gepner models. A hypersurface family corresponding to a hypersurface $\widetilde{W}=0$ which defines a supervariety \mathcal{M} embedded in $\mathbf{WSP}(n_{z_1},\ldots,n_{z_{n+1}}|n_{\eta_1},\ldots,n_{\eta_{2m}})$ is denoted by $\mathbf{WSP}(n_{z_1},\ldots,n_{z_{n+1}}|n_{\eta_1},\ldots,n_{\eta_{2m}})[d]$. Here, \widetilde{W} is the modified superpotential obtained by satisfying (2.8) and d is its degree of quasihomogeneity. The fermionic weights are

determined by requiring (1.8) and (1.9) to be satisfied. The solutions for these fermionic weights are parameterized by $k=1,\ldots, [\frac{d}{2s_k}]$ and $l=1,\ldots, [\frac{d}{2s_l}]$, where s_k and s_l are the coefficients of k and l in the first and third fermion weight assignments, respectively. For each hypersurface family, the Hodge numbers $h^{1,1}$ and $h^{2,1}$ and the Euler number $\chi=2\left(h^{1,1}-h^{2,1}\right)$ of the associated Landau-Ginzburg orbifold are given. When the Newton polytope of \widetilde{W}_{bos} admits a nef partition, this is indicated by "nef". In a number of cases, the Landau-Ginzburg orbifold can be given a geometrical interpretation as a product of a two-torus and a K3 surface, which is indicated by " $T^2 \times K3$ ". Finally, when the Newton polytope of \widetilde{W}_{bos} is nonreflexive Gorenstein, this is indicated by "nonRG".

#	hypersurface family	$h^{1,1}$	$h^{2,1}$	χ	Comments
1	$\mathbf{WSP}(1, 1, 1, 1, 1, 1, 1, 2 k, 4 - k)[4]$	0	90	-180	
2	$\mathbf{WSP}(1, 1, 1, 2, 2, 2, 3 k, 6 - k)[6]$	0	84	-168	
3	WSP(2,3,1,2,2,2,4 2k,8-2k)[8]	1	73	-144	nonRG
4	$\mathbf{WSP}(2,3,2,3,1,1,4 k,8-k)[8]$	1	77	-152	
5	$\mathbf{WSP}(2,4,2,4,1,2,5 2k,10-2k)[10]$	1	85	-168	
6	$\mathbf{WSP}(2,3,3,3,3,4,6 3k,12-3k)[12]$	1	61	-120	
7	$\mathbf{WSP}(2,2,3,3,4,4,6 2k,12-2k)[12]$	3	51	-96	
8	$\mathbf{WSP}(1,3,3,3,4,4,6 k,12-k)[12]$	10	46	-72	
9	$\mathbf{WSP}(1, 2, 3, 4, 4, 4, 6 2k, 12 - 2k)[12]$	2	62	-120	
10	$\mathbf{WSP}(1, 1, 4, 4, 4, 4, 6 2k, 12 - 2k)[12]$	7	79	-144	
11	$\mathbf{WSP}(2,5,2,3,3,3,6 k,12-k)[12]$	9	39	-60	
12	WSP(2, 5, 2, 2, 3, 4, 6 2k, 12 - 2k)[12]	2	74	-144	nonRG
13	WSP(2,5,1,3,3,4,6 k,12-k)[12]	1	61	-120	nonRG
14	WSP(2,5,1,2,4,4,6 2k,12-2k)[12]	3	75	-144	nonRG
15	$\mathbf{WSP}(2,5,2,5,2,2,6 2k,12-2k)[12]$	2	128	-252	
16	$\mathbf{WSP}(2,5,2,5,1,3,6 k,12-k)[12]$	3	69	-132	
17	$\mathbf{WSP}(4, 6, 2, 7, 1, 4, 8 2k, 16 - 2k)[16]$	3	75	-144	
18	WSP(2,7,2,7,2,4,8 2k,16-2k)[16]	4	148	-288	nef
19	$\mathbf{WSP}(6, 4, 2, 3, 6, 6, 9 6, 12)[18]$	2	62	-120	nonRG
20	$\mathbf{WSP}(6, 4, 6, 4, 1, 6, 9 2k, 18 - 2k)[18]$	2	56	-108	
21	WSP(1, 2, 6, 6, 6, 6, 9 6, 12)[18]	8	68	-120	
22	WSP(2, 8, 2, 3, 6, 6, 9 2k, 18 - 2k)[18]	8	68	-120	
23	$\mathbf{WSP}(6, 4, 2, 8, 1, 6, 9 2k, 18 - 2k)[18]$	4	76	-144	
24	WSP(2, 8, 2, 8, 1, 6, 9 2k, 18 - 2k)[18]	2	110	-216	r m2 - 120
25	$\mathbf{WSP}(4, 8, 4, 4, 5, 5, 10 2k, 20 - 2k)[20]$	21	21	0	nef, $T^2 \times K3$
26	WSP(4, 8, 4, 8, 1, 5, 10 2k, 20 - 2k)[20]	13	49	-72	
27	WSP(2, 9, 4, 5, 5, 5, 10 k, 20-k)[20]	17	29	-24	D.C.
28	WSP(4, 8, 2, 9, 2, 5, 10 2k, 20 - 2k)[20]	7	79	-144	nonRG
29	WSP(2, 9, 2, 9, 4, 4, 10 2k, 20 - 2k)[20]	7	143	-272	
30	WSP(3, 3, 6, 8, 8, 8, 12 2k, 24 - 2k)[24]	21	21	0	ner, $I = \times K3$
31	WSP(1,3,8,8,8,8,12 4k,24-4k)[24]	16	52	-72	
32	WSP(6, 9, 3, 4, 6, 8, 12 6k, 24 - 6k)[24]	3	51	-96	D.C
33 34	WSP $(6, 9, 1, 6, 6, 8, 12 6k, 24 - 6k)[24]$ WSP $(6, 9, 2, 3, 8, 8, 12 2k, 24 - 2k)[24]$	10 13	46 37	-72 -48	on RG $ on RG$
35	WSP $(6, 9, 2, 3, 8, 8, 12 2k, 24 - 2k)[24]$ WSP $(6, 9, 1, 4, 8, 8, 12 2k, 24 - 2k)[24]$	12	48	-48 -72	nonRG
36	WSP $(6, 9, 6, 9, 2, 4, 12 2k, 24 - 2k)[24]$	6	66	-120	nef
37	WSP ($6, 9, 6, 9, 2, 4, 12 0k, 24 - 0k)[24]$ WSP ($4, 10, 3, 3, 8, 8, 12 2k, 24 - 2k)[24]$	11	35	-120 -48	nei
38	WSP (6, 9, 4, 10, 3, 4, 12 2 k , 24 - 2 k)[24]	11	35	-48	
39	WSP $(6, 9, 4, 10, 3, 4, 12 2k, 24 - 2k)[24]$ WSP $(6, 9, 4, 10, 1, 6, 12 2k, 24 - 2k)[24]$	7	55	-48 -96	nonRG
40	WSP $(2, 11, 3, 6, 6, 8, 12 2k, 24 - 2k)[24]$	10	46	-30 -72	пошто
41	WSP (2,11,3,4,8,8,12 2k,24-2k)[24]	12	48	-72	nonRG
42	WSP (2,11,1,6,8,8,12 2k,24-2k)[24]	9	81	-144	пошто
43	WSP $(6, 9, 2, 11, 4, 4, 12 2k, 24 - 2k)[24]$	6	90	-168	nef
44	WSP $(6, 9, 2, 11, 2, 4, 12 2k, 24 - 2k)[24]$	6	114	-216	nef
45	WSP $(4, 10, 2, 11, 2, 6, 12 2k, 24 - 2k)[24]$	7	55	-96	nonRG
46	WSP (4, 10, 2, 11, 1, 8, $12 2k, 24 - 2k)[24]$ WSP (4, 10, 2, 11, 1, 8, $12 2k, 24 - 2k)[24]$	3	99	-192	пошто
40	VV SF (4, 10, 2, 11, 1, 0, 12 2K, 24 - 2K)[24]	3	99	-192	

47	$\mathbf{WSP}(2,11,2,11,4,6,12 2k,24-2k)[24]$	8	164	-312	nef
48	$\mathbf{WSP}(2,11,2,11,2,8,12 2k,24-2k)[24]$	3	243	-480	nef
49	WSP(4, 12, 2, 13, 4, 7, 14 2k, 28 - 2k)[28]	8	80	-144	_
50	$\mathbf{WSP}(3, 6, 6, 10, 10, 10, 15 2k, 30 - 2k)[30]$	21	21	0	nef, $T^2 \times K3$
51	WSP(2, 3, 10, 10, 10, 10, 15 10, 20)[30]	17	41	-48	
52	$\mathbf{WSP}(6, 12, 5, 6, 6, 10, 15 6k, 30 - 6k)[30]$	21	21	0	nef, $T^2 \times K3$
53	$\mathbf{WSP}(6, 12, 2, 5, 10, 10, 15 2k, 30 - 2k)[30]$	17	41	-48	
54	$\mathbf{WSP}(6, 12, 1, 6, 10, 10, 15 2k, 30 - 2k)[30]$	7	55	-96	nonRG
55	$\mathbf{WSP}(2, 14, 3, 6, 10, 10, 15 2k, 30 - 2k)[30]$	7	55	-96	
56	$\mathbf{WSP}(6, 12, 2, 14, 5, 6, 15 2k, 30 - 2k)[30]$	7	55	-96	nonRG
57	$\mathbf{WSP}(6, 12, 2, 14, 1, 10, 15 2k, 30 - 2k)[30]$	5	101	-192	
58	$\mathbf{WSP}(2, 14, 2, 14, 3, 10, 15 2k, 30 - 2k)[30]$	5	101	-192	0
59	$\mathbf{WSP}(12, 8, 4, 9, 9, 12, 18 6k, 36 - 6k)[36]$	21	21	0	nef, $T^2 \times K3$
60	$\mathbf{WSP}(12, 8, 1, 9, 12, 12, 18 6k, 36 - 6k)[36]$	10	46	-72	
61	$\mathbf{WSP}(12, 8, 6, 15, 4, 9, 18 6k, 36 - 6k)[36]$	8	44	-72	nonRG
62	$\mathbf{WSP}(12, 8, 6, 15, 1, 12, 18 6k, 36 - 6k)[36]$	13	49	-72	nonRG
63	$\mathbf{WSP}(4, 16, 4, 9, 9, 12, 18 2k, 36 - 2k)[36]$	21	21	0	nef, $T^2 \times K3$
64	$\mathbf{WSP}(4, 16, 1, 9, 12, 12, 18 2k, 36 - 2k)[36]$	20	56	-72	
65	$\mathbf{WSP}(6, 15, 4, 16, 4, 9, 18 2k, 36 - 2k)[36]$	14	50	-72	
66	$\mathbf{WSP}(6, 15, 4, 16, 1, 12, 18 2k, 36 - 2k)[36]$	5	77	-144	
67	$\mathbf{WSP}(12, 8, 2, 17, 6, 9, 18 2k, 36 - 2k)[36]$	11	53	-84	
68	$\mathbf{WSP}(12, 8, 2, 17, 3, 12, 18 2k, 36 - 2k)[36]$	13	49	-72	nonRG
69	$\mathbf{WSP}(2, 17, 2, 9, 12, 12, 18 2k, 36 - 2k)[36]$	16	100	-168	
70	$\mathbf{WSP}(6, 15, 2, 17, 2, 12, 18 2k, 36 - 2k)[36]$	5	185	-360	nef
71	$\mathbf{WSP}(4, 16, 2, 17, 6, 9, 18 2k, 36 - 2k)[36]$	13	73	-120	nonRG
72	WSP(4, 16, 2, 17, 3, 12, 18 2k, 36 - 2k)[36]	5	77	-144	nonRG
73	$\mathbf{WSP}(2, 17, 2, 17, 4, 12, 18 2k, 36 - 2k)[36]$	7	271	-528	nef
74	$\mathbf{WSP}(10, 15, 10, 15, 2, 8, 20 10k, 40 - 10k)[40]$	7	63	-112	nef
75	$\mathbf{WSP}(10, 15, 8, 16, 1, 10, 20 2k, 40 - 2k)[40]$	13	49	-72	
76	$\mathbf{WSP}(10, 15, 4, 18, 5, 8, 20 2k, 40 - 2k)[40]$	19	27	-16	
77	$\mathbf{WSP}(10, 15, 2, 19, 4, 10, 20 2k, 40 - 2k)[40]$	12	96	-168	nef
78	WSP(8, 16, 2, 19, 5, 10, 20 2k, 40 - 2k)[40]	13	49	-72	c
79	$\mathbf{WSP}(2, 19, 2, 19, 8, 10, 20 2k, 40 - 2k)[40]$	11	227	-432	nef
80	WSP(1, 6, 14, 14, 14, 14, 21 14, 28)[42]	23	47	-48	
81	WSP(6, 18, 6, 18, 1, 14, 21 6k, 42 - 6k)[42]	15	63	-96	
82	WSP(2, 20, 6, 7, 14, 14, 21 2k, 42 - 2k)[42]	23	47	-48	
83	WSP(6, 18, 2, 20, 3, 14, 21 2k, 42 - 2k)[42]	15	63	-96	
84	WSP(6, 21, 1, 12, 16, 16, 24 2k, 48 - 2k)[48]	20	56	-72	
85	WSP(8, 20, 6, 21, 1, 16, 24 2k, 48 - 2k)[48]	8	68	-120	c
86	WSP(6, 21, 6, 21, 2, 16, 24 6k, 48 - 6k)[48]	7	127	-240	nef
87	WSP $(6, 21, 4, 22, 3, 16, 24 2k, 48 - 2k)[48]$ WSP $(2, 23, 3, 12, 16, 16, 24 2k, 48 - 2k)[48]$	17	41	-48	
88		20	56	-72	
89 90	$\mathbf{WSP}(12, 18, 2, 23, 1, 16, 24 2k, 48 - 2k)[48]$ $\mathbf{WSP}(8, 20, 2, 23, 3, 16, 24 2k, 48 - 2k)[48]$	9 8	129 68	-240 -120	nonRG
91	WSP $(6, 21, 2, 23, 8, 10, 24 2k, 48 - 2k)[48]$	16	112	-120	Homney
92	WSP $(6, 21, 2, 23, 6, 12, 24 2k, 46 - 2k)[48]$ WSP $(6, 21, 2, 23, 4, 16, 24 2k, 48 - 2k)[48]$	8	164	-312	nonRG
93	WSP $(0, 21, 2, 23, 4, 10, 24 2k, 46 - 2k)[48]$	9	321	-624	nef
94	WSP (3, 12, 15, 20, 20, 20, 30 10 k , 60 – 10 k)[60]	21	21	0	nef, $T^2 \times K3$
95	WSP $(3, 12, 16, 20, 20, 20, 30 16k, 60 - 16k)[60]$	21	21	0	nef, $T^2 \times K3$
96	WSP(10, 25, 3, 12, 20, 20, 30 10k, 60 - 10k)[60]	23	23	0	nei, 1 × 110
97	WSP(12, 24, 10, 25, 4, 15, 30 2k, 60 - 2k)[60]	29	29	0	
98	WSP(6, 27, 10, 12, 15, 20, 30 6k, 60 - 6k)[60]	15	39	-48	
99	WSP $(6, 27, 5, 12, 10, 20, 30 0k, 60 - 6k)[60]$	23	23	0	
100	WSP(6, 27, 2, 15, 20, 20, 30 2k, 60 - 2k)[60]	31	55	-48	
101	WSP(12, 24, 6, 27, 1, 20, 30 6k, 60 - 6k)[60]	17	65	-96	
102	WSP (10, 25, 6, 27, 10, 12, 30 2 k , 60 – 2 k)[60]	13	49	-72	
103	WSP(10, 25, 6, 27, 2, 20, 30 2k, 60 - 2k)[60]	10	106	-192	nonRG
104	WSP(6, 27, 6, 27, 4, 20, 30 6k, 60 - 6k)[60]	11	107	-192	nef
105	WSP(4, 28, 3, 15, 20, 20, 30 2k, 60 - 2k)[60]	31	31	0	
106	WSP(10, 25, 4, 28, 3, 20, 30 2k, 60 - 2k)[60]	10	46	-72	nonRG
107	WSP(6, 27, 4, 28, 10, 15, 30 2k, 60 - 2k)[60]	25	37	-24	nonRG

108	WSP(6, 27, 4, 28, 5, 20, 30 2k, 60 - 2k)[60]	10	46	-72	nonRG
109	WSP(2, 29, 12, 12, 15, 20, 30 2k, 60 - 2k)[60]	23	47	-48	nomva
110	WSP(2, 29, 4, 15, 20, 20, 30 2k, 60 - 2k)[60]	26	86	-120	
111	WSP(12, 24, 2, 29, 3, 20, 30 2k, 60 - 2k)[60]	17	65	-96	
112	WSP(10, 25, 2, 29, 12, 12, 30 2k, 60 - 2k)[60]	25	85	-120	
113	WSP(10, 25, 2, 29, 4, 20, 30 2k, 60 - 2k)[60]	11	155	-188	nef
114	WSP(6, 27, 2, 29, 6, 20, 30 2k, 60 - 2k)[60]	10	178	-336	
115	WSP(4, 28, 2, 29, 12, 15, 30 2k, 60 - 2k)[60]	17	101	-168	
116	WSP $(18, 27, 4, 34, 1, 24, 36 2k, 72 - 2k)[72]$	14	98	-168	
117	WSP $(18, 27, 2, 35, 8, 18, 36 2k, 72 - 2k)[72]$	19	91	-144	nef
118	WSP(18, 27, 2, 35, 2, 24, 36 2k, 72 - 2k)[72]	14	242	-456	
119	WSP(6, 33, 2, 35, 8, 24, 36 2k, 72 - 2k)[72]	15	183	-336	
120	WSP(4, 34, 2, 35, 9, 24, 36 2k, 72 - 2k)[72]	14	98	-168	nonRG
121	WSP(12, 36, 1, 21, 28, 28, 42 2k, 84 - 2k)[84]	45	45	0	
122	WSP(14, 35, 12, 36, 1, 28, 42 2k, 84 - 2k)[84]	15	63	-96	
123	WSP(6, 39, 4, 21, 28, 28, 42 2k, 84 - 2k)[84]	41	41	0	
124	WSP(14, 35, 6, 39, 4, 28, 42 2k, 84 - 2k)[84]	16	76	-120	nonRG
125	WSP(4, 40, 12, 21, 21, 28, 42 2k, 84 - 2k)[84]	21	21	0	nef, $T^2 \times K3$
126	WSP(14, 35, 4, 40, 12, 21, 42 2k, 84 - 2k)[84]	35	35	0	
127	WSP(2, 41, 6, 21, 28, 28, 42 2k, 84 - 2k)[84]	40	76	-72	
128	WSP(14, 35, 2, 41, 6, 28, 42 2k, 84 - 2k)[84]	15	147	-264	nef
129	WSP(12, 36, 2, 41, 14, 21, 42 2k, 84 - 2k)[84]	34	58	-48	
130	WSP(12, 36, 2, 41, 7, 28, 42 2k, 84 - 2k)[84]	15	63	-96	
131	WSP(2, 41, 2, 41, 12, 28, 42 2k, 84 - 2k)[84]	11	491	-960	nef
132	WSP(30, 20, 18, 36, 1, 30, 45 6k, 90 - 6k)[90]	29	41	-24	nonRG
133	$\mathbf{WSP}(18, 36, 10, 40, 1, 30, 45 2k, 90 - 2k)[90]$	17	65	-96	
134	$\mathbf{WSP}(30, 20, 2, 44, 9, 30, 45 2k, 90 - 2k)[90]$	29	41	-24	nonRG
135	$\mathbf{WSP}(18, 36, 2, 44, 5, 30, 45 2k, 90 - 2k)[90]$	17	65	-96	
136	$\mathbf{WSP}(10, 40, 2, 44, 9, 30, 45 2k, 90 - 2k)[90]$	17	65	-96	
137	$\mathbf{WSP}(24, 36, 6, 45, 1, 32, 48 6k, 96 - 6k)[96]$	24	84	-120	
138	$\mathbf{WSP}(24, 36, 2, 47, 3, 32, 48 2k, 96 - 2k)[96]$	24	84	-120	
139	$\mathbf{WSP}(6, 45, 2, 47, 12, 32, 48 2k, 96 - 2k)[96]$	18	222	-408	
140	$\mathbf{WSP}(30, 45, 1, 24, 40, 40, 60 10k, 120 - 10k)[120]$	39	39	0	nonRG
141	WSP(24, 48, 10, 55, 3, 40, 60 2k, 120 - 2k)[120]	29	29	0	
142	$\mathbf{WSP}(30, 45, 8, 56, 1, 40, 60 2k, 120 - 2k)[120]$	24	84	-120	
143	$\mathbf{WSP}(30, 45, 6, 57, 2, 40, 60 6k, 120 - 6k)[120]$	23	143	-240	
144	$\mathbf{WSP}(24, 48, 6, 57, 5, 40, 60 6k, 120 - 6k)[120]$	29	29	0	
145	$\mathbf{WSP}(10, 55, 6, 57, 12, 40, 60 2k, 120 - 2k)[120]$	22	82	-120	
146	$\mathbf{WSP}(30, 45, 4, 58, 3, 40, 60 2k, 120 - 2k)[120]$	29	53	-48	
147	$\mathbf{WSP}(6,57,4,58,15,40,60 2k,120-2k)[120]$	29	53	-48	nonRG
148	$\mathbf{WSP}(2,59,15,24,40,40,60 2k,120-2k)[120]$	39	39	0	
149	$\mathbf{WSP}(30, 45, 2, 59, 20, 24, 60 2k, 120 - 2k)[120]$	33	69	-72	
150	$\mathbf{WSP}(30, 45, 2, 59, 4, 40, 60 2k, 120 - 2k)[120]$	24	204	-360	
151	$\mathbf{WSP}(10, 55, 2, 59, 24, 30, 60 2k, 120 - 2k)[120]$	33	141	-216	
152	$\mathbf{WSP}(8, 56, 2, 59, 15, 40, 60 2k, 120 - 2k)[120]$	24	84	-120	
153	$\mathbf{WSP}(10, 65, 4, 68, 28, 35, 70 2k, 140 - 2k)[140]$	47	47	0	
154	WSP(28, 56, 2, 69, 20, 35, 70 2k, 140 - 2k)[140]	53	53	0	
155	WSP(2,77,12,39,52,52,78 2k,156-2k)[156]	71	71	0	c
156	WSP(26, 65, 2, 77, 12, 52, 78 2k, 156 - 2k)[156]	23	143	-240	nef
157	WSP(42, 63, 12, 78, 1, 56, 84 6k, 168 - 6k)[168]	38	74	-72	
158	WSP(42, 63, 8, 80, 3, 56, 84 2k, 168 – 2k)[168]	39	39	0	
159	WSP(42, 63, 6, 81, 4, 56, 84 6k, 168 – 6k)[168]	33	105	-144	
160	WSP $(8, 80, 6, 81, 21, 56, 84 2k, 168 - 2k)[168]$ WSP $(42, 63, 2, 83, 6, 56, 84 2k, 168 - 2k)[168]$	39	39	919	
161		34	190	-312	
162	WSP ($12, 78, 2, 83, 21, 56, 84 2k, 168 - 2k)[168]$	38	74	-72 -624	
163 164	WSP $(6, 81, 2, 83, 24, 56, 84 2k, 168 - 2k)[168]$ WSP $(18, 81, 2, 89, 20, 60, 90 2k, 180 - 2k)[180]$	23 42	$\frac{335}{150}$	-024 -216	
165	WSP (18, 81, 2, 89, 20, 00, $90 2k$, $180 - 2k$)[180] WSP (54, 81, 2, 107, 8, 72, $108 2k$, $216 - 2k$)[216]	42	180	-216 -264	
166	WSP (54, 81, 2, 107, 8, 72, 108 2 k , 210 - 2 k)[210] WSP (20, 100, 2, 109, 44, 55, 110 2 k , 220 - 2 k)[220]	48 71	71	-264 0	
167	WSP (20, 100, 2, 109, 44, 55, 110 2 k , 220 – 2 k)[220] WSP (48, 96, 30, 105, 1, 80, 120 6 k , 240 – 6 k)[240]	53	53	0	
168	WSP (48, 96, 30, 103, 1, 80, 120 0 k , 240 – 6 k)[240] WSP (48, 96, 2, 119, 15, 80, 120 2 k , 240 – 2 k)[240]	53	53	0	
100	(40, 30, 2, 113, 10, 00, 120 2h, 240 - 2h)[240]	99	99	U	

169	WSP(30, 105, 2, 119, 24, 80, 120 2k, 240 - 2k)[240]	50	134	-168	
170	$\mathbf{WSP}(66, 99, 6, 129, 8, 88, 132 6k, 264 - 6k)[264]$	57	81	-48	
171	WSP(78, 117, 24, 144, 1, 104, 156 6k, 312 - 6k)[312]	69	69	0	
172	WSP(78, 117, 2, 155, 12, 104, 156 2k, 312 - 2k)[312]	66	174	-216	
173	WSP(24, 144, 2, 155, 39, 104, 156 2k, 312 - 2k)[312]	69	69	0	
174	WSP(14, 161, 2, 167, 48, 112, 168 2k, 336 - 2k)[336]	47	287	-480	
175	WSP(14, 203, 6, 207, 60, 140, 210 2k, 420 - 2k)[420]	59	131	-144	
176	WSP(150, 225, 2, 299, 24, 200, 300 2k, 600 - 2k)[600]	119	167	-96	
177	WSP(42, 441, 2, 461, 132, 308, 462 2k, 924 - 2k)[924]	137	257	-240	
178	$\mathbf{WSP}(1, 2, 1, 2, 1, 2, 1 k, 5 - k)[5]$	1	85	-168	
179	WSP(1, 1, 2, 2, 2, 2, 2 2, 4)[6]	1	73	-144	
180	WSP(2,3,2,3,2,2,2 2k,8-2k)[8]	2	86	-168	nef
181	$\mathbf{WSP}(2,3,2,3,2,3,1 k,8-k)[8]$	2	58	-112	
182	WSP(3, 2, 3, 2, 3, 2, 3 k, 9 - k)[9]	8	35	-54	nef
183	WSP(3, 2, 1, 3, 3, 3, 3 3, 6)[9]	2	62	-120	nonRG
184	WSP(3, 2, 3, 2, 1, 4, 3 k, 9 - k)[9]	2	56	-108	
185	WSP(1,4,1,3,3,3,3 k,9-k)[9]	8	68	-120	
186	WSP(3, 2, 1, 4, 1, 4, 3 k, 9 - k)[9]	4	76	-144	
187	$\mathbf{WSP}(1,4,1,4,1,4,3 k,9-k)[9]$	2	110	-216	
188	WSP(3, 3, 3, 3, 4, 4, 4 k, 12 - k)[12]	21	21	0	nef, $T^2 \times K3$
189	WSP(2, 3, 3, 4, 4, 4, 4 2k, 12 - 2k)[12]	11	35	-48	
190	WSP(1, 3, 4, 4, 4, 4, 4 4, 8)[12]	2	62	-120	
191	$\mathbf{WSP}(2,5,3,3,3,4,4 k,12-k)[12]$	4	40	-72	
192	$\mathbf{WSP}(2,5,2,3,4,4,4 2k,12-2k)[12]$	2	62	-120	nonRG
193	$\mathbf{WSP}(2,5,1,4,4,4,4 2k,12-2k)[12]$	9	57	-96	
194	$\mathbf{WSP}(2, 5, 2, 5, 3, 3, 4 k, 12 - k)[12]$	5	41	-72	
195	$\mathbf{WSP}(2, 5, 2, 5, 2, 4, 4 2k, 12 - 2k)[12]$	5	101	-192	nef
196	WSP(2, 5, 2, 5, 2, 5, 3 k, 12 - k)[12]	3	57	-108	2
197	$\mathbf{WSP}(3,6,3,3,5,5,5 k,15-k)[15]$	21	21	0	nef, $T^2 \times K3$
198	$\mathbf{WSP}(3,6,1,5,5,5,5 k,15-k)[15]$	17	41	-48	
199	$\mathbf{WSP}(3,6,1,7,3,5,5 k,15-k)[15]$	7	55	-96	nonRG
200	$\mathbf{WSP}(3,6,1,7,1,7,5 k,15-k)[15]$	5	101	-192	
201	WSP(4, 6, 2, 7, 2, 7, 4 2k, 16 - 2k)[16]	8	104	-192	nef
202	$\mathbf{WSP}(4, 8, 4, 8, 2, 9, 5 2k, 20 - 2k)[20]$	13	49	-72	
203	$\mathbf{WSP}(1, 10, 3, 7, 7, 7, 7 k, 21 - k)[21]$	23	47	-48	
204	WSP(3, 9, 3, 9, 1, 10, 7 k, 21 - k)[21]	15	63	-96	nef, $T^2 \times K3$
205	WSP(6, 9, 3, 6, 8, 8, 8 2k, 24 - 2k)[24]	21	21	0	ner, $T^2 \times K3$
206	WSP(6, 9, 1, 8, 8, 8, 8 2k, 24 - 2k)[24]	24	36	-24	of
207	WSP(6, 9, 6, 9, 4, 6, 8 6k, 24 - 6k)[24]	7	55 42	-96	nef
208 209	$\mathbf{WSP}(6, 9, 6, 9, 2, 8, 8 2k, 24 - 2k)[24]$ $\mathbf{WSP}(6, 9, 4, 10, 3, 8, 8 2k, 24 - 2k)[24]$	19 9	43	-48	nef
210	WSP (6, 9, 4, 10, 5, 8, 8 2k, 24 - 2k)[24] WSP (6, 9, 6, 9, 4, 10, 4 2k, 24 - 2k)[24]	20	33 32	-48 -24	of
210	WSP $(0, 9, 0, 9, 4, 10, 4 2k, 24 - 2k)[24]$ WSP $(2, 11, 3, 8, 8, 8, 8 2k, 24 - 2k)[24]$	24	36	-24 -24	nef
212	WSP (6, 9, 2, 11, 6, 6, 8 2k, 24 - 2k)[24]	8	68	-120	nonRG
213	WSP $(6, 9, 2, 11, 4, 8, 8 2k, 24 - 2k)[24]$	10	70	-120	nonRG
214	WSP $(6, 9, 4, 10, 2, 11, 6 2k, 24 - 2k)[24]$	10	70	-120	nonRG
215	WSP(2, 11, 2, 11, 6, 8, 8 2k, 24 - 2k)[24]	11	131	-240	nef
216	WSP(4, 10, 2, 11, 2, 11, 8 2k, 24 - 2k)[24]	9	153	-288	nef
217	$\mathbf{WSP}(12, 8, 2, 17, 9, 12, 12 2k, 36 - 2k)[36]$	10	46	-72	1101
218	$\mathbf{WSP}(12, 8, 6, 15, 2, 17, 12 2k, 36 - 2k)[36]$ $\mathbf{WSP}(12, 8, 6, 15, 2, 17, 12 2k, 36 - 2k)[36]$	17	77	-120	nonRG
219	WSP(4, 16, 2, 17, 9, 12, 12 2k, 36 - 2k)[36]	20	56	-72	nomed
220	WSP(6, 15, 4, 16, 2, 17, 12 2k, 36 - 2k)[36]	13	109	-192	nef
221	WSP (10, 15, 10, 15, 4, 18, $8 2k$, $40 - 2k$)[40]	31	23	16	nef
222	WSP (10, 15, 8, 16, 2, 19, $10 2k$, $40 - 2k$)[40]	23	59	-72	
223	WSP (15, 10, 9, 18, 1, 22, 15 k , 45 - k)[45]	29	41	-24	nonRG
224	WSP (9, 18, 5, 20, 1, 22, 15 k , 45 - k)[45]	17	65	-96	
225	WSP(6, 21, 6, 21, 4, 22, 16 2k, 48 - 2k)[48]	31	55	-48	nef
226	WSP(6, 21, 2, 23, 12, 16, 16 2k, 48 - 2k)[48]	26	86	-120	
227	WSP(8, 20, 6, 21, 2, 23, 16 2k, 48 - 2k)[48]	18	102	-168	
228	WSP(12, 18, 2, 23, 2, 23, 16 2k, 48 - 2k)[48]	13	229	-432	nef
229	WSP(6, 27, 12, 15, 20, 20, 20 2k, 60 - 2k)[60]	21	21	0	nef, $T^2 \times K3$
•	• • • • • • • • • • • • • • • • • • • •				

230	$\mathbf{WSP}(10, 25, 6, 27, 12, 20, 20 2k, 60 - 2k)[60]$	31	31	0	
231	WSP(6, 27, 4, 28, 15, 20, 20 2k, 60 - 2k)[60]	31	31	0	
232	$\mathbf{WSP}(10, 25, 6, 27, 4, 28, 20 2k, 60 - 2k)[60]$	22	58	-72	nonRG
233	WSP(12, 24, 6, 27, 2, 29, 20 2k, 60 - 2k)[60]	21	117	-192	
234	WSP(18, 27, 4, 34, 2, 35, 24 2k, 72 - 2k)[72]	22	130	-216	
235	WSP(12, 36, 2, 41, 21, 28, 28 2k, 84 - 2k)[84]	45	45	0	
236	WSP(14, 35, 12, 36, 2, 41, 28 2k, 84 - 2k)[84]	37	85	-96	
237	WSP(24, 36, 6, 45, 2, 47, 32 2k, 96 - 2k)[96]	26	158	-264	
238	WSP(24, 48, 10, 55, 6, 57, 40 2k, 120 - 2k)[120]	49	49	0	
239	WSP(30, 45, 6, 57, 4, 58, 40 2k, 120 - 2k)[120]	43	67	-48	
240	WSP(30, 45, 2, 59, 24, 40, 40 2k, 120 - 2k)[120]	55	55	0	
241	WSP(30, 45, 8, 56, 2, 59, 40 2k, 120 - 2k)[120]	34	118	-168	
242	WSP(42, 63, 8, 80, 6, 81, 56 2k, 168 - 2k)[168]	55	55	0	
243	WSP(42, 63, 12, 78, 2, 83, 56 2k, 168 - 2k)[168]	46	106	-120	nonRG
244	WSP(48, 96, 30, 105, 2, 119, 80 2k, 240 - 2k)[240]	89	89	0	
245	WSP(78, 117, 24, 144, 2, 155, 104 2k, 312 - 2k)[312]	97	97	0	
246	WSP(1, 2, 2, 2, 2, 2, 2, 2, 3 2, 4, 2, 4)[6]	0	84	-168	
247	WSP(2, 3, 2, 3, 2, 3, 2, 3, 4 k, 8 - k, 2, 6)[8]	1	53	-104	
	WSP(2, 3, 2, 3, 2, 3, 2, 3, 4 3, 5, 4, 4)[8]	1	53	-104	
248	WSP(3, 3, 4, 4, 4, 4, 4, 4, 6 2k, 12 - 2k, 4, 8)[12]	21	21	0	$T^2 \times K3$
249	WSP(2, 5, 3, 4, 4, 4, 4, 4, 6 2k, 12 - 2k, 4, 8)[12]	2	62	-120	nonRG
250	WSP(2, 5, 2, 5, 4, 4, 4, 4, 6 2k, 12 - 2k, 2l, 12 - 2l)[12]	7	79	-144	
251	$\mathbf{WSP}(6, 9, 6, 9, 6, 8, 8, 8, 12 2k, 24 - 2k, 6l, 24 - 6l)[24]$	21	21	0	nef, $T^2 \times K3$
252	WSP(6, 9, 6, 9, 4, 10, 8, 8, 12 2k, 24 - 2k, 6l, 24 - 6l)[24]	11	35	-48	
253	WSP(6, 9, 2, 11, 8, 8, 8, 8, 12 2k, 24 - 2k, 4l, 24 - 4l)[24]	16	52	-72	
254	WSP(1, 1, 1, 1, 1, 1, 1, 1, 1, 1, 1, 1, 2, 1, 2)[3]	0	84	-168	

Table 2: Supervariety hypersurface families associated with $\hat{c}=3$ Gepner models.

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